

# Perspectives on Nodal Superconductors

Kazumi Maki,<sup>1</sup> Stephan Haas,<sup>1</sup> David Parker,<sup>1</sup> and Hyekyung Won<sup>2</sup>

<sup>1</sup>*Department of Physics and Astronomy,*

*University of Southern California, Los Angeles, CA 90089-0484 USA*

<sup>2</sup>*Department of Physics, Hallym University, Chuncheon 200-702, South Korea*

(Dated: February 2, 2008)

## Abstract

In the last few years the gap symmetries of many new superconductors, including  $\text{Sr}_2\text{RuO}_4$ ,  $\text{CeCoIn}_5$ ,  $\kappa\text{-(ET)}_2\text{Cu(NCS)}_2$ ,  $\text{YNi}_2\text{B}_2\text{C}$  and  $\text{PrOs}_4\text{Sb}_{12}$ , have been identified via angle-dependent magnetothermal conductivity measurements. However, a controversy still persists as to the nature of the superconductivity in  $\text{Sr}_2\text{RuO}_4$ . For  $\text{PrOs}_4\text{Sb}_{12}$ , spin-triplet superconductivity has recently been proposed. Here, we also propose g-wave superconductivity for  $\text{UPd}_2\text{Al}_3$  (i.e.,  $\Delta(\mathbf{k}) = \Delta \cos(2\chi)$ ,  $\chi = ck_z$ ) based on recent thermal conductivity data.

PACS numbers:

## 1. Introduction

After the appearance of heavy-fermion superconductors and organic superconductors in 1979 the gap symmetries of these new compounds have been a central issue[1]. However, until recently only the  $d_{x^2-y^2}$ -wave symmetry of the gap function  $\Delta(\mathbf{k})$  in high- $T_c$  cuprates has been established by the elegant Josephson interferometry[2] and the angle resolved photoemission spectra (ARPES)[3]. Unfortunately, so far these powerful techniques are unavailable for heavy-fermion superconductors and organic superconductors with lower superconducting transition temperatures  $T_c \leq 10K$ .

In the last few years, Izawa et al have established the gap symmetries of superconductivity in  $\text{Sr}_2\text{RuO}_4$ [4],  $\text{CeCoIn}_5$ [5],  $\kappa\text{-ET}_2\text{Cu}(\text{NCS})_2$ [6],  $\text{YNi}_2\text{B}_2\text{C}$ [7] and  $\text{PrOs}_4\text{Sb}_{12}$ [8] through the angle dependent magnetothermal conductivity. This breakthrough relies in part on the availability of high-quality single crystals of these compounds and in part on the theoretical development initiated by Volovik.[9] Last year, we have reviewed the progress in [10].

In the present paper, we focus on 3 recent topics in nodal superconductors. In spite of ample evidence for f-wave superconductivity in  $\text{Sr}_2\text{RuO}_4$  [10] the controversy regarding this compound appears to continue. Therefore in section 2 we discuss the angle dependent magnetospecific heat data by Deguchi et al [11]. Now evidence for spin-triplet superconductivity in  $\text{PrOs}_4\text{Sb}_{12}$  is mounting. In section 3, we describe p+h-wave superconductivity for the A and B phases in  $\text{PrOs}_4\text{Sb}_{12}$ [12]. Recently angle-dependent thermal conductivity data in the vortex state in  $\text{UPd}_2\text{Al}_3$  has been reported.[13] In section 4 we analyze the angle-dependent magnetothermal conductivity  $\kappa_{yy}$  when the field is rotated within the z-x plane, and we conclude that  $\Delta(\mathbf{k})$  in  $\text{UPd}_2\text{Al}_3$  is given by  $\Delta(\mathbf{k}) = \Delta \cos(2\chi)$ [14]. In Fig. 1 we show the new  $|\Delta(\mathbf{k})|$ 's so far identified.

## 2. F-wave Superconductivity in $\text{Sr}_2\text{RuO}_4$

Superconductivity in  $\text{Sr}_2\text{RuO}_4$  was discovered in 1994[15].  $\text{Sr}_2\text{RuO}_4$  is an isocrystal to  $\text{La}_2\text{CuO}_4$ , but it is metallic down to low temperatures and becomes superconducting around  $T = 1.5$  K. An early review on  $\text{Sr}_2\text{RuO}_4$  can be found in Ref.[16]. From the analogy to superfluid  $^3\text{He}$  Rice and Sigrist[17] proposed 2D p-wave superconductivity. Indeed spin-triplet pairing and related chiral symmetry-breaking have been established [18, 19, 20]. As sample quality improved around 1999, both the specific heat data [21] and the superfluid density [22] indicated nodal structure in the superconducting order parameter of  $\text{Sr}_2\text{RuO}_4$ . These findings ruled out p-wave superconductivity and its generalization[23]. Therefore, a variety

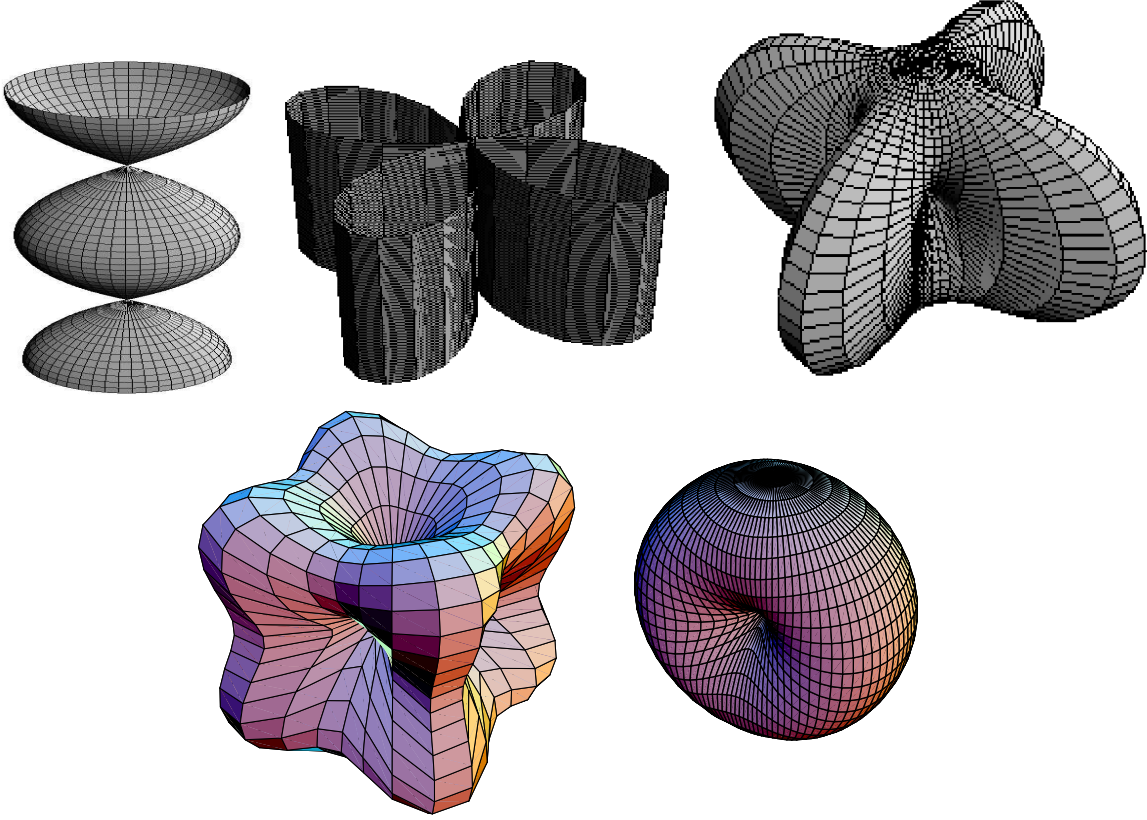


FIG. 1: From top left, 2D f-wave -  $\text{Sr}_2\text{RuO}_4$ ,  $d_{x^2-y^2}$ -wave -  $\text{CeCoIn}_5$  and  $\kappa\text{-(ET)}_2\text{Cu(NCS)}_2$  s+g-wave -  $\text{YNi}_2\text{B}_2\text{C}$ , p+h-wave -  $\text{PrOs}_4\text{Sb}_{12}$  - A phase, p+h-wave -  $\text{PrOs}_4\text{Sb}_{12}$  - B phase.

of f-wave order parameters were suggested. [24] In Fig.2 and Fig.3 we show the specific heat data [21] and the superfluid density data[22] compared with a variety of models.

However, these experiments cannot tell us about the nodal structure of  $\Delta(\mathbf{k})$ . In a quasi-2D system such as  $\text{Sr}_2\text{RuO}_4$ , the line nodes in  $\Delta(\mathbf{k})$  can be either vertical or horizontal. But vertical nodes are incompatible with the angular dependent magnetothermal conductivity [4] and the ultrasonic attenuation data [25]. Furthermore, Ref. 4 indicates that the horizontal nodes are far away from  $\chi_0=0$ . This suggests  $\Delta(\mathbf{k}) = \mathbf{d}e^{\pm i\phi} \cos(\chi)$ , i.e. 2D f-wave superconductivity[26].

This interpretation is contested by Deguchi et al [11]. They measured the magnetospecific heat of  $\text{Sr}_2\text{RuO}_4$  in a rotating magnetic field down to 100 mK and found cusp-like features only in the regime  $0.12K < T < 0.31K$ . From our earlier analysis of s+g-wave superconductivity[27, 28], we deduce that there should be a point-like minigap with

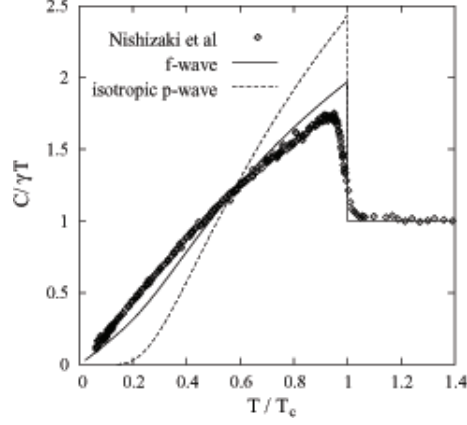


FIG. 2: Specific heat for 2D p-wave and f-wave models for  $\text{Sr}_2\text{RuO}_4$ .

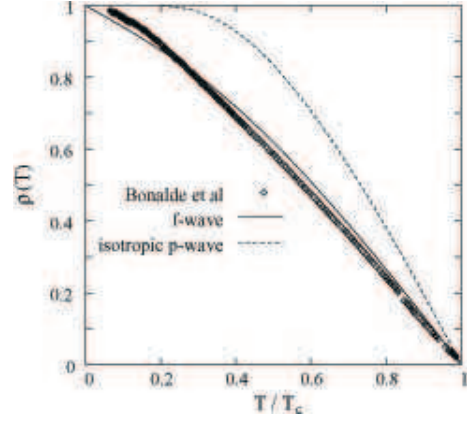


FIG. 3: Superfluid density for 2D p-wave and f-wave models for  $\text{Sr}_2\text{RuO}_4$ .

$\Delta_{min} \sim 0.1K$ . The simplest triplet gap function which has these minigaps is

$$\Delta(\mathbf{k}) = \mathbf{d}e^{\pm i\phi}(1 + a \cos(4\phi) \cos(\chi)) \quad (1)$$

where  $|1 - a| \leq 0.1$ . Deguchi et al have proposed the Miyake-Nariyiko (MN) model [29], in order to describe the measured specific heat. However, it is easy to see that the MN model cannot give the cusp-like features in the magnetospecific heat. Also, the MN model cannot describe the observed  $T^2$  specific heat or the T-linear dependence of the superfluid density. Moreover, the angular dependent thermal conductivity data and the universal heat conduction in  $\kappa_{xx}$  by Suzuki et al [30] are incompatible with the MN model. Therefore further experiments on  $\text{Sr}_2\text{RuO}_4$  are highly desirable. We have proposed that the optical conductivity [31], the Raman scattering[32] and the supercurrent [33, 34] in  $\text{Sr}_2\text{RuO}_4$  will provide further insight on its superconductivity.

### 3. Triplet Superconductivity in PrOs<sub>4</sub>Sb<sub>12</sub>

Superconductivity with  $T_c = 1.8$  K has been discovered very recently in the skutterudite PrOs<sub>4</sub>Sb<sub>12</sub>[35, 36, 37]. Angle-dependent thermal conductivity data on this system has revealed a multi-phase structure, characterized by a gap function with point nodes.[8] In order to account for this nodal structure s+g-wave superconductivity has been proposed. [10, 38]

Recently there has been mounting experimental evidence for triplet superconductivity in this compound. First, from  $\mu$ SR measurements Aoki et al discovered a remnant magnetization in the B-phase of this compound, indicating triplet pairing. [39] Second, the thermal conductivity measurement in a magnetic field down to low-temperature ( $T > 150$  mK) indicates  $\kappa_{zz} \sim T$  and  $H$  [40], consistent with triplet pairing. Later we shall discuss  $\kappa_{zz}$  measured in a magnetic field rotated within the z-x plane. This data is fully consistent with triplet p+h-wave superconductivity in PrOs<sub>4</sub>Sb<sub>12</sub>. Finally, a recently reported NMR result for the Knight shift by Tou et al [41] also suggests the triplet pairing. Here we propose p+h-wave order parameters

$$\Delta_A(\mathbf{k}) = \frac{3}{2} \mathbf{d} e^{\pm i\phi_1 \pm i\phi_2 \pm i\phi_3} (1 - \hat{k}_1^4 - \hat{k}_2^4 - \hat{k}_3^4) \quad (2)$$

$$\Delta_B(\mathbf{k}) = \mathbf{d} e^{\pm i\phi_3} (1 - \hat{k}_3^4) \quad (3)$$

for the A-phase and B-phase of PrOs<sub>4</sub>Sb<sub>12</sub>, respectively, where  $e^{\pm i\phi_1} = \hat{k}_2 \pm i\hat{k}_3$ , etc. These order parameters have nodal structures consistent with the angle dependent thermal conductivity data [8], assuming that in the experiment the nodes in the B-phase are aligned parallel to the y-axis. We note  $|\Delta(\mathbf{k})|$  in the A-phase has cubic symmetry whereas in the B-phase it has axial symmetry. Furthermore, it appears that in the slow field-cooled situation the nodes in the B-phase are aligned parallel to the magnetic field. At least this is the simplest way to interpret the superfluid density measurement by Chia et al [42, 43].

Here we give expressions for the thermal conductivity  $\kappa_{zz}$  in a magnetic field in the superclean limit ( $\sqrt{\Gamma\Delta} \ll v\sqrt{eH}$ ),

$$\kappa_{zz}/\kappa_n = \frac{v^2 e H}{8\Delta^2} \sin^2(\theta) \quad , A - phase \quad (4)$$

$$= \frac{3v^2 e H}{64\Delta^2} \sin^2(\theta) \quad , B - phase \quad (5)$$

and in the clean limit ( $v\sqrt{eH} \ll \sqrt{\Gamma\Delta}$ )

$$\kappa_{zz}/\kappa_{00} = 1 + \frac{3v^2eH}{40\Gamma\Delta} \ln\left(\sqrt{\frac{2\Delta}{\Gamma}}\right) \sin^2(\theta) \ln\left(\frac{\Delta}{v\sqrt{eH}\sin(\theta)}\right), \quad A-phase \quad (6)$$

$$= 1 + \frac{v^2eH}{12\Gamma\Delta} \ln\left(\sqrt{\frac{2\Delta}{\Gamma}}\right) \sin^2(\theta) \ln\left(\frac{\Delta}{v\sqrt{eH}\sin(\theta)}\right), \quad B-phase \quad (7)$$

where  $\kappa_n$  and  $\kappa_{00}$  are the thermal conductivity in the normal state and the thermal conductivity in the limit of universal heat conduction  $\Gamma \rightarrow 0, T \rightarrow 0$ . Here,  $\Gamma$  is the quasiparticle scattering rate in the normal state, and  $\theta$  is the angle  $\mathbf{H}$  makes from the  $\hat{z}$  axis. In both Eq.(5) and Eq.(7) we have assumed that the nodes in the B-phase are parallel to the z axis. Otherwise  $\kappa_{zz}$  is smaller by a factor of  $10 \sim 50$ . In Fig. 4 we compare the observed angle dependent thermal conductivity with Eq.(6) and (7). These equations give an excellent fit.

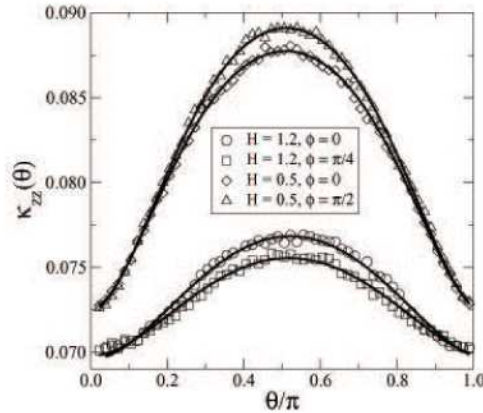


FIG. 4: Angular-dependent thermal conductivity in  $\text{PrOs}_4\text{Sb}_{12}$ .

From this we extract  $v = 0.96 \times 10^7$  cm/sec and  $\Gamma = 0.1$  K, where use is made of the weak-coupling theory gaps  $\Delta_A = 4.2K$  and  $\Delta_B = 3.5K$  for the A and B phase respectively. Note that de Haas-van Alphen measurements [44] give comparable values of  $v$  ( $0.7 \times 10^7$  cm/sec [ $\alpha$ -band],  $0.6 \times 10^7$  cm/sec [ $\beta$ -band] and  $0.23 \times 10^7$  cm/sec [ $\gamma$  band]).

#### 4. G-wave superconductivity in $\text{UPd}_2\text{Al}_3$

This heavy-fermion superconductor with  $T_c \simeq 2K$  was discovered by Geibel et al [45] in 1991. The reduction of the Knight shift in the superconducting state seen in NMR [46] and the Pauli limiting of  $H_{c2}$  in  $\text{UPd}_2\text{Al}_3$ [47] established spin-singlet pairing. Nodal superconductivity with horizontal nodes has been deduced from the thermal conductivity

[48] and the c-axis tunneling data from UPd<sub>2</sub>Al<sub>3</sub> thin films [49]. Here we focus on thermal conductivity data reported in Ref.[13]. First,  $\kappa_{zz}$  at T = 0.4 K in a rotating magnetic field was measured. For H < 0.5 T no  $\phi$  dependence was seen, ( $\phi$  is the angle  $\mathbf{H}$  makes from the x axis). This indicates that the nodal lines should be horizontal. Second,  $\kappa_{yy}$  in a magnetic field rotated within the z-x plane was measured.

We show this in Fig. 5. Following the standard procedure [10], the thermal conductivity

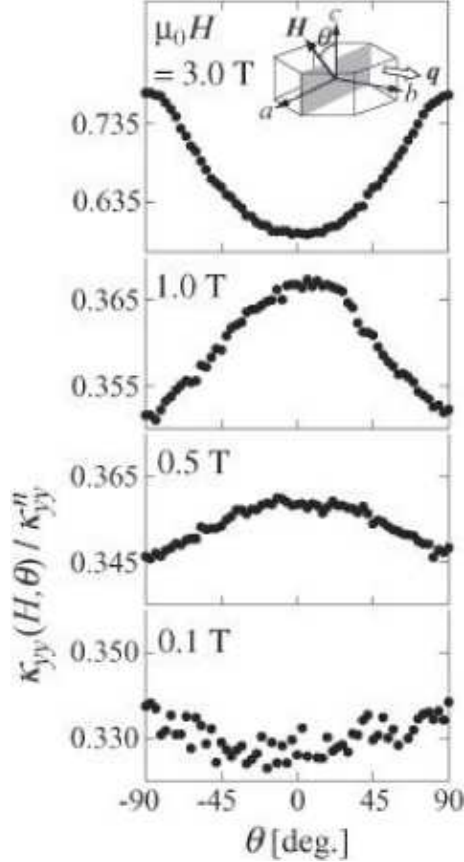


FIG. 5: Angular dependent magnetothermal conductivity in UPd<sub>2</sub>Al<sub>3</sub>

$\kappa_{yy}$  is obtained for a variety of  $\Delta(\mathbf{k})$  with horizontal nodes as

$$\kappa_{yy}/\kappa_n = \frac{2}{\pi} \frac{v_a^2 e H}{\Delta^2} F_1(\theta) \quad (8)$$

in the superclean limit and

$$\kappa_{yy}/\kappa_{00} = 1 + \frac{v_a^2 e H}{6\pi\Gamma\Delta} F_2(\theta) \ln\left(2\sqrt{\frac{2\Delta}{\pi\Gamma}}\right) \ln\left(\frac{\Delta}{v\sqrt{eH}}\right) \quad (9)$$

in the clean limit, where

$$F_1(\theta) = \sqrt{\cos^2 \theta + \alpha \sin^2 \theta} (1 + \sin^2 \theta (-\frac{3}{8} + \alpha \sin^2 \chi_0)) \quad (10)$$

$$F_2(\theta) = \sqrt{\cos^2 \theta + \alpha \sin^2 \theta} (1 + \sin^2 \theta (-\frac{1}{4} + \alpha \sin^2 \chi_0)), \quad (11)$$

and  $\alpha = (v_c/v_a)^2$ .  $\theta$  is the angle  $\mathbf{H}$  makes from the z-axis and  $\chi_0$  is the nodal position. For  $\Delta(\mathbf{k}) \sim \cos \chi$ ,  $\cos(2\chi)$ , and  $\sin \chi$  we obtain  $\chi_0 = \frac{\pi}{2}$ ,  $\frac{\pi}{4}$  and 0, respectively. We show  $F_1(\theta)$  and  $F_2(\theta)$  in Fig. 6 where we used  $\alpha = 0.69$ , the appropriate value for UPd<sub>2</sub>Al<sub>3</sub>. A comparison of Fig. 5 and Fig. 6 indicates that  $\Delta(\mathbf{k}) \sim \cos(2\chi)$  is the most appropriate choice. Similarly

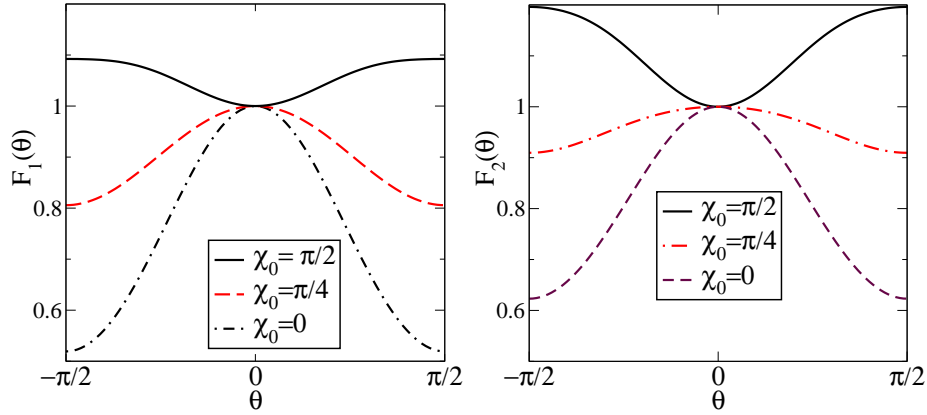


FIG. 6: Angular functions  $F_1(\theta)$  (left) and  $F_2(\theta)$  for various nodal positions

the universal heat conduction in nodal superconductors [50, 51] for a variety of quasi-2D systems ( $f = \cos(2\phi), \sin(2\phi), \cos \chi, \cos(2\chi), \sin \chi$ ) is a quantity of interest. We obtain [51]

$$\frac{\kappa_{xx}}{\kappa_n} = \frac{2\Gamma_0}{\pi\Delta} \frac{1}{\sqrt{1+C_0^2}} E\left(\frac{1}{\sqrt{1+C_0^2}}\right) = I_1\left(\frac{\Gamma}{\Gamma_0}\right) \quad (12)$$

for all  $f$ 's given above. Here  $\Gamma_0 = 0.866T_c$  and  $C_0$  is determined by

$$\frac{C_0^2}{\sqrt{1+C_0^2}} K\left(\frac{1}{\sqrt{1+C_0^2}}\right) = \frac{\pi\Gamma}{2\Delta}, \quad (13)$$

and  $K(z)$  and  $E(z)$  are the complete elliptic integrals. Here  $\kappa_n$  is the thermal conductivity in the normal state with  $\Gamma = \Gamma_0$ . Eq.(12) tells us that  $\kappa_{xx}$  cannot discriminate between different nodal structures. On the other hand, we find

$$\frac{\kappa_{zz}}{\kappa_n} = I_1\left(\frac{\Gamma}{\Gamma_0}\right) \text{ for } f = \cos(2\phi), \sin(2\phi), \cos(2\chi), \quad (14)$$



but

$$\frac{\kappa_{zz}}{\kappa_n} = \frac{4\Gamma_0}{\pi\Delta\sqrt{1+C_0^2}} \left( E\left(\frac{1}{\sqrt{1+C_0^2}}\right) - C_0^2 \left( K\left(\frac{1}{\sqrt{1+C_0^2}}\right) - E\left(\frac{1}{\sqrt{1+C_0^2}}\right) \right) \right) \quad (15)$$

$$= I_2\left(\frac{\Gamma}{\Gamma_0}\right) \text{ for } f = \cos\chi, e^{\pm i\phi} \cos\chi \quad (16)$$

and

$$\frac{\kappa_{zz}}{\kappa_n} = \frac{4\Gamma_0 C_0^2}{\pi\Delta\sqrt{1+C_0^2}} \left( K\left(\frac{1}{\sqrt{1+C_0^2}}\right) - E\left(\frac{1}{\sqrt{1+C_0^2}}\right) \right) \quad (17)$$

$$= I_3\left(\frac{\Gamma}{\Gamma_0}\right) \quad (18)$$

for  $f=\sin\chi$  and  $e^{\pm i\phi} \sin\chi$ . We show  $I_1(\Gamma/\Gamma_0)$ ,  $I_2(\Gamma/\Gamma_0)$  and  $I_3(\Gamma/\Gamma_0)$  in Fig. 7. Watanabe et

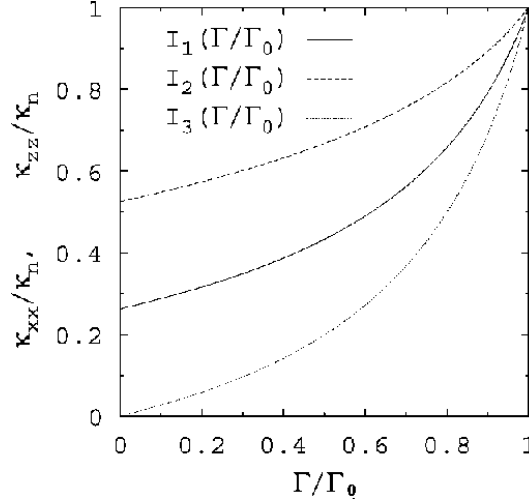


FIG. 7: The functions  $I_1$ ,  $I_2$  and  $I_3$

al [13] also measured  $\kappa_{xx}$  and  $\kappa_{zz}$  as a function of  $\mathbf{H}(\|\hat{z})$ . Of course the effects of impurity scattering and of magnetic fields are very different. Nevertheless the comparison of these figures suggests again  $\Delta(\mathbf{k}) \sim \cos(2\chi)$ .

## 5. Concluding Remarks

We have surveyed recent developments on nodal superconductors. As to the superconductivity in  $\text{Sr}_2\text{RuO}_4$  we believe the 2-D f-wave model with horizontal nodes is most promising, in spite of the new specific heat data by Deguchi et al. However, further experiments on  $\text{Sr}_2\text{RuO}_4$  are clearly desirable.

The p+h-wave superconductivity in  $\text{PrOs}_4\text{Sb}_{12}$  appears to solve many jigsaw puzzles simultaneously. These superconducting order parameters are highly degenerate due to mul-

triple chiral symmetry breaking. We expect exciting topological defects in these systems associated with the chiral symmetry breaking.

Also UPd<sub>2</sub>Al<sub>3</sub> is the first Uranium compound examined through angle-dependent thermal conductivity experiments. From a more indirect way, the gap symmetry of UPt<sub>3</sub> has been deduced to be  $f = e^{\pm 2i\phi} \sin^2 \theta \cos \theta$  or E<sub>2u</sub> at least for the B phase[52]. It has been shown that there are many triplet superconductors, including UPt<sub>3</sub>, UBe<sub>13</sub>, URu<sub>2</sub>Si<sub>2</sub> and UNi<sub>2</sub>Al<sub>3</sub> [53]. The determination of the gap symmetries of these superconductors is of great interest.

Of course the gap symmetry itself cannot tell the underlying pairing mechanism of these systems. But at least this provides the first important step for further exploration. Also, phonons most likely play no role in the pairing mechanism of most nodal superconductors. The majority of these pairings appear to be due to the antiparamagnon exchange. But we can expect more exotic interactions as well in this plethora of nodal superconductors.

### Acknowledgments

We have benefitted from helpful collaborations and discussions with Balazs Dora, Koichi Izawa, Hae-Young Kee, Yuji Matsuda, Peter Thalmeier, Attila Virosztek and Tadataka Watanabe. K.M. acknowledges gratefully the hospitality of the Max Planck Institute of the Physics of Complex Systems at Dresden, where a part of this work was done. S.H. was supported by the NSF, Grant No. DMR-0089882.

- 
- [1] M. Sigrist and K. Ueda, Rev. Mod. Phys. **63**, 239 (1991).
  - [2] C. C. Tsuei and J. R. Kirtley, Rev. Mod. Phys. **72**, 969 (2000).
  - [3] A. Damascelli, Z. Hussain and Z.X. Shen, Rev. Mod. Phys. **75**, 473 (2003).
  - [4] K. Izawa et al, Phys. Rev. Lett. **86**, 2653 (2001).
  - [5] K. Izawa et al, Phys. Rev. Lett. **87**, 57002 (2001).
  - [6] K. Izawa et al, Phys. Rev. Lett. **88**, 27002 (2002).
  - [7] K. Izawa et al, Phys. Rev. Lett. **89**, 137006 (2002).
  - [8] K. Izawa et al, Phys. Rev. Lett. **90**, 11701 (2003).
  - [9] G.E. Volovik, JETP Letters **58**, 469 (1993).
  - [10] H. Won, Q. Yuan, P. Thalmeier and K. Maki, Brazilian J. Phys. **33**, 675 (2003).
  - [11] K. Deguchi, Z.Q. Mao, H. Yaguchi and Y. Maeno, Phys. Rev. Lett. **92**, 047002 (2004); K.

- Deguchi, Z.Q. Mao, Y. Maeno, J. Phys. Soc. Jpn. **73**, 1313 (2004).
- [12] K. Maki, S. Haas, D. Parker, H. Won, K. Izawa and Y. Matsuda, cond-mat/0406492.
  - [13] T. Watanabe et al, cond-mat/0405211.
  - [14] H. Won and K. Maki, SCES 2004 proceeding.
  - [15] Y. Maeno et al, Nature **372**, 532 (1994).
  - [16] A.P. Mackenzie and Y. Maeno, Rev. Mod. Phys. **75**, 657 (2003).
  - [17] T.M. Rice and M. Sigrist, J. Phys. Cond. Matt. **7**, L643 (1995).
  - [18] G.M. Luke et al, Nature **394**, 558 (1998).
  - [19] K. Ishida et al, Nature **396**, 653 (1998).
  - [20] R. Jin, Y. Liu, Z.Q. Mao and Y. Maeno, Europhys. Lett **51**, 341 (2000).
  - [21] S. Nishizaki, Y. Maeno and Z.Q. Mao, J. Phys. Soc. Jpn. **69**, 572 (2001).
  - [22] I. Bonalde et al, Phys. Rev. Lett. **85**, 4775 (2000).
  - [23] M.E. Zhitomirsky and T.M. Rice, Phys. Rev. Lett. **87**, 057001 (2001).
  - [24] T. Dahm, H. Won and K. Maki, cond-mat/0006307.
  - [25] C. Lupien et al, Phys. Rev. Lett. **86**, 5986 (2001).
  - [26] K. Maki and H. Won in “Fluctuating Paths and Fields”, edited by W. Jahnke et al, World Scientific, Singapore 2001.
  - [27] K. Maki, P. Thalmeier and H. Won, Phys. Rev. **B65**, R140502 (2002).
  - [28] K. Maki, H. Won and S. Haas, Phys. Rev. **B69**, 012502 (2004).
  - [29] K. Miyake and O. Narikiyo, Phys. Rev. Lett. **83**, 1423 (1999).
  - [30] M. Suzuki et al, Phys. Rev. Lett. **88**, 227004 (2002).
  - [31] B. Dora, K. Maki, A. Virosztek, Europhys. Lett, **62**, 426 (2003).
  - [32] H-Y. Kee, K. Maki, C.H. Chung, Phys. Rev. **B67**, 186504(R) (2003).
  - [33] H-Y. Kee, Y.B. Kim and K. Maki, Phys. Rev. B (in press).
  - [34] I. Khavkine, H-Y. Kee and K. Maki, cond-mat/0405236.
  - [35] E.D. Bauer, N.A. Frederick, P-C. Ho, V.S. Zapf and M.B Maple, Phys. Rev **B65**, R100506 (2002).
  - [36] R. Vollmer, M. Etzkorn, P.S. Anil Kumer, H. Ibach and J. Kirschner, Phys. Rev. Lett. **90**, 57001 (2003).
  - [37] H. Kotegawa, M. Yogi, Y. Imamura, Y. Kawasaki, G.-q. Zheng, Y. Kitaoka, S. Ohsaki, H. Sugawara, Y. Aoki, H. Sato, Phys. Rev. Lett. **90**, 27001 (2003).

- [38] K. Maki, H. Won, P. Thalmeier, Q. Yuan, K. Izawa and Y. Matsuda, Europhys. Lett. **64**, 496 (2003).
- [39] K. Aoki et al, Phys. Rev. Lett. **91**, 067003 (2003).
- [40] K. Izawa et al (unpublished).
- [41] H. Tou (private communication).
- [42] E.E. Chia, M.B. Salomon, H. Sugawara, and H. Sato, Phys. Rev. Lett. **91**, 247003 (2003).
- [43] H. Won, D. Parker, S. Haas and K. Maki, Current Appl. Phys. **4**, 523 (2004).
- [44] H. Sugawara, S. Osaki, S.R. Saha, Y. Aoki, H. Sato, Y. Inada, H. Shishido, R. Settai, Y. Onuki, H. Harima and K. Oikawa, Phys. Rev. B **66** 220504(R), (2002).
- [45] C. Geibel et al, Z. Phys. B **84**, 1 (1991).
- [46] H. Tou et al, J. Phys. Soc. Jpn. **64**, 725 (1995).
- [47] J. Hessert et al, Physica B **230-232**, 373 (1997).
- [48] May Chiao, B. Lussier, E. Elleman and L. Taillefer, Physica B **230-232**, 370 (1997).
- [49] M. Jourdan, M. Huth and H. Adrian, Nature **398**, 47 (1999).
- [50] P.A. Lee, Phys. Rev. Lett. **71**, 1887 (1993).
- [51] Y. Sun and K. Maki, Europhys. Lett. **32**, 335 (1995).
- [52] P. Thalmeier and K. Maki, Phys. Rev. B **67**, 92510 (2003).
- [53] H. Tou, K. Ishida, Y. Kitaoka, cond-mat/0308562.